

## Assignment 13

Due date: Wednesday, May 10

### *Relativistic string equation of motion*

Show that the Euler-Lagrange equations<sup>1</sup>,

$$0 = \partial_p \left( \frac{\partial \mathcal{L}}{\partial (\partial_p s^\alpha)} \right) + \partial_q \left( \frac{\partial \mathcal{L}}{\partial (\partial_q s^\alpha)} \right) = \partial_a \left( \frac{\partial \mathcal{L}}{\partial (\partial_a s^\alpha)} \right),$$

applied to the Lagrangian of the relativistic string produces the equation of motion

$$0 = \epsilon^{ab} (\partial_a v^{\alpha\beta}) (\partial_b s_\beta), \quad (1)$$

where

$$v^{\alpha\beta} = a^{\alpha\beta} / \mathcal{L} \quad (2)$$

is a normalized rescaling of the anti-symmetric tangent-space tensor  $a^{\alpha\beta}$ , analogous to the 4-velocity we define for the point particle.

### *Relativistic string as a non-linear modification of simple elastic string*

Consider a relativistic string that resides only in the  $(x, y)$  plane and has infinite extent in  $x$ . The simple (non-relativistic) elastic string, in this geometry, is described by a function  $y(x, t)$  that gives the displacement as a function of time  $t$  at all positions  $x$ . Suppose we likewise use  $(x, t)$  as the parameters of the relativistic string's world-surface:

$$s^\alpha(x, t) = (ct, x, y(x, t), 0).$$

The action of a relativistic string with this parameterization is a functional of  $y(x, t)$ :

$$S[y] = \int \mathcal{L}(\partial_x y, \partial_t y) dx dt.$$

Find the Lagrangian function  $\mathcal{L}$  and from that show that the resulting (Euler-Lagrange) equation of motion for  $y$  is the wave equation for the simple elastic string with additional cubic terms. Show that arbitrary running waves,  $y = f(x - ct)$  and  $y = g(x + ct)$ , are solutions, but that this does not extend to linear combinations of right and left moving waves.

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<sup>1</sup>Einstein summation convention for all repeated indices, greek and latin.

*Local conservation law for relativistic string*

Show that the stress-energy tensor density for relativistic strings,

$$T^{\alpha\beta}(x) = \int (\mathcal{L} dp dq) v^{\alpha\gamma} v_{\gamma}^{\beta} \delta^4(x - s(p, q)),$$

is locally conserved, that is,

$$\partial_{\beta} T^{\alpha\beta}(x) = 0.$$

Here  $v^{\alpha\beta}$  is the normalized tensor (2). Your derivation should mirror the derivation in lecture of the conservation law for the point-particle stress-energy tensor. In particular, you will need to use the string's equations of motion (1).